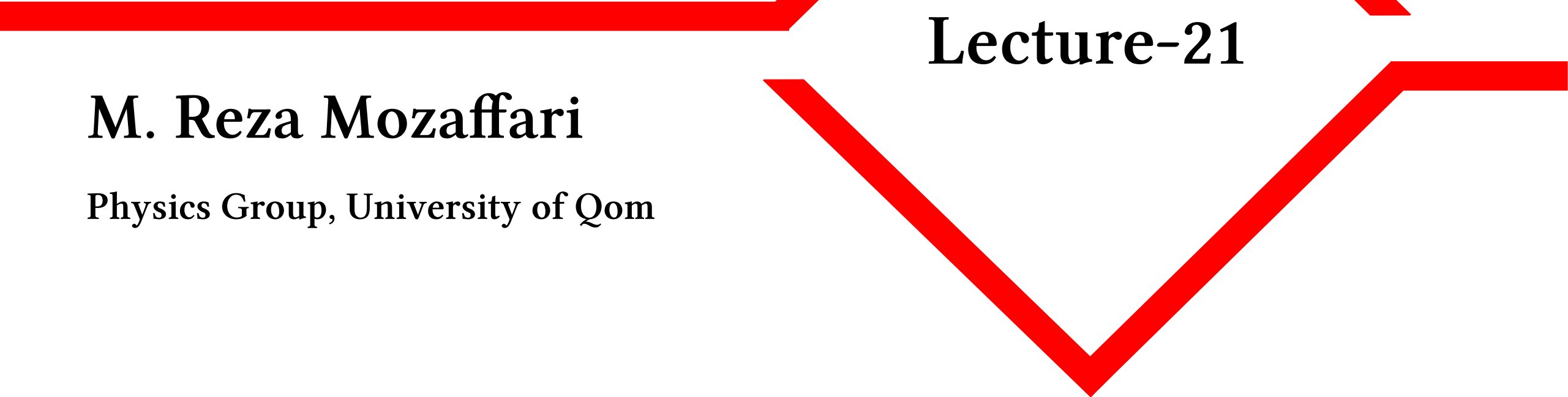


Computational Physics



Lecture-01

M. Reza Mozaffari

Physics Group, University of Qom

Contents

- Basis Concepts

Basic Concepts

- Computational Physics aims at **solving physical problems** by means of **numerical methods developed** in the field of numerical analysis

- In most cases it is **not possible to find analytic solutions** and one must rely on good approximations. $\int_a^b e^{-x^2} dx$

- In contrast to the much simpler integral $\int_a^b e^x dx = e^a - e^b$

- Hence, we have to approximate this integral in such a way that the approximation is accurate enough for our purpose. $\int_a^b e^{-x^2} dx \approx \dots$

Basic Concepts

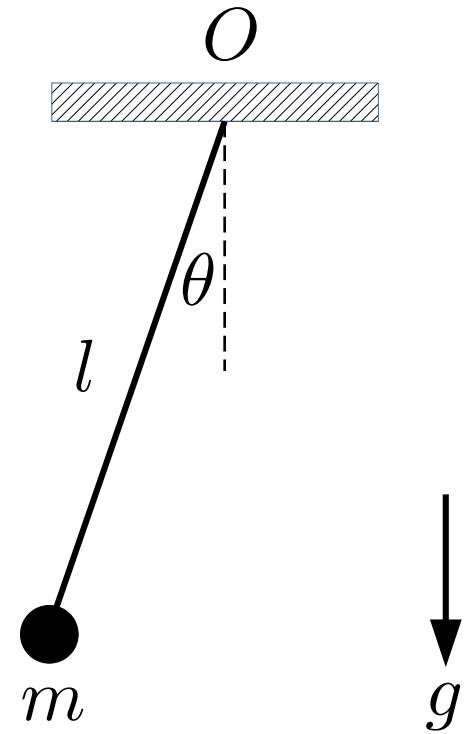
- We know from basic mechanics that the time evolution of a frictionless pendulum of mass m and length l in a gravitational field is modeled by the differential equation

$$\ddot{\theta} + \frac{g}{l} \sin \theta = 0$$

$$\theta(t = 0) = \theta_0 \text{ and } \dot{\theta}(t = 0) = 0$$

$$\theta \ll 1 : \sin \theta \approx \theta \quad \ddot{\theta} + \frac{g}{l} \theta = 0$$

$$\theta = \theta_0 \cos(\omega t), \quad \omega = \sqrt{\frac{g}{l}}, \quad T_0 = 2\pi \sqrt{\frac{l}{g}}$$



Basic Concepts

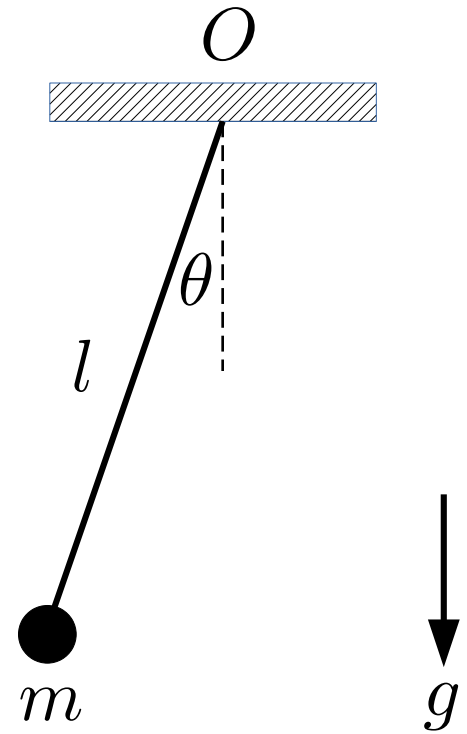
- We know from basic mechanics that the time evolution of a frictionless pendulum of mass m and length l in a gravitational field is modeled by the differential equation

$$\ddot{\theta} + \frac{g}{l} \sin \theta = 0$$

$$\theta(t = 0) = \theta_0 \text{ and } \dot{\theta}(t = 0) = 0$$

$$\ddot{\theta} = \frac{d\dot{\theta}}{dt} = \frac{d\dot{\theta}}{d\theta} \frac{d\theta}{dt} = \dot{\theta} \frac{d\dot{\theta}}{d\theta}$$

$$\dot{\theta} \frac{d\dot{\theta}}{d\theta} + \frac{g}{l} \sin \theta = 0 \Rightarrow \dot{\theta} d\dot{\theta} = -\frac{g}{l} \sin \theta d\theta \Rightarrow \int_0^{\dot{\theta}} \dot{\theta} d\dot{\theta} = -\frac{g}{l} \int_{\theta_0}^{\theta} \sin \theta d\theta$$



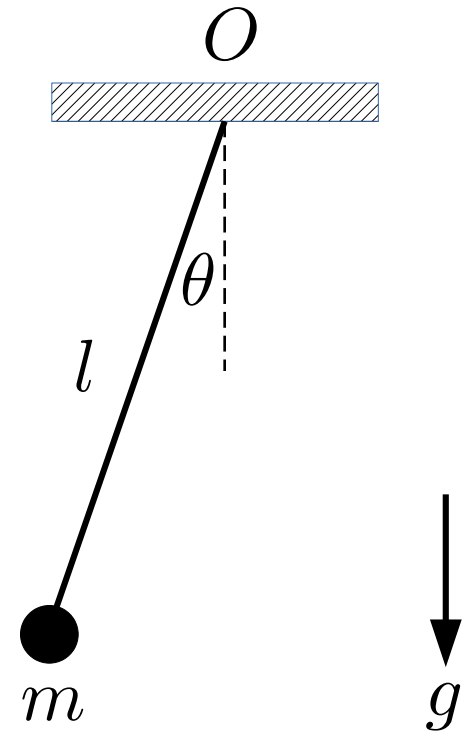
Basic Concepts

- We know from basic mechanics that the time evolution of a frictionless pendulum of mass m and length l in a gravitational field is modeled by the differential equation

$$\int_0^{\dot{\theta}} \dot{\theta} d\dot{\theta} = -\frac{g}{l} \int_{\theta_0}^{\theta} \sin \theta d\theta$$

$$\frac{1}{2} \dot{\theta}^2 = \frac{g}{l} (\cos \theta - \cos \theta_0) = \frac{g}{l} \left(1 - 2 \sin^2 \frac{\theta}{2} - 1 + 2 \sin^2 \frac{\theta_0}{2} \right)$$

$$\frac{1}{2} \dot{\theta}^2 = 2 \frac{g}{l} \left(\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2} \right) \Rightarrow \dot{\theta} = 2 \sqrt{\frac{g}{l}} \sqrt{\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2}}$$



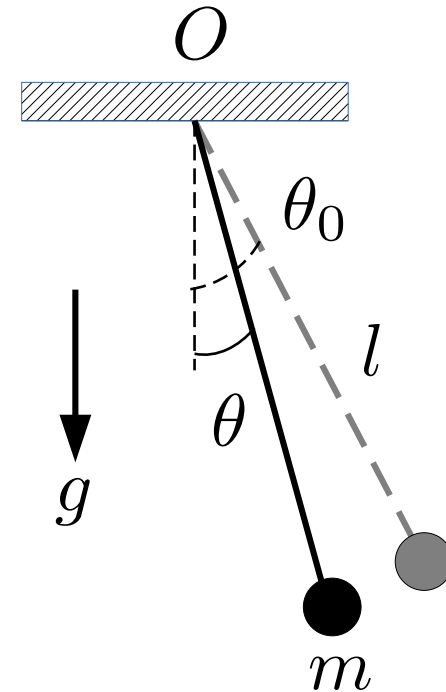
Basic Concepts

- We know from basic mechanics that the time evolution of a frictionless pendulum of mass m and length l in a gravitational field is modeled by the differential equation

$$\dot{\theta} = 2\sqrt{\frac{g}{l}} \sqrt{\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2}}$$

$$\frac{d\theta}{dt} = 2\sqrt{\frac{g}{l}} \sqrt{\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2}} \Rightarrow dt = \frac{1}{2} \sqrt{\frac{l}{g}} \frac{d\theta}{\sqrt{\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2}}}$$

$$\frac{1}{4} \int_0^T dt = \frac{1}{2} \sqrt{\frac{l}{g}} \int_0^{\theta_0} \frac{d\theta}{\sqrt{\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2}}}$$

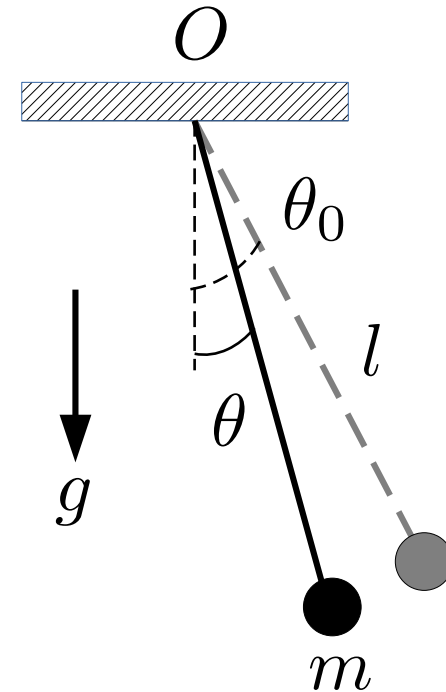


Basic Concepts

- A frictionless pendulum

$$T = 2\sqrt{\frac{l}{g}} \int_0^{\theta_0} \frac{d\theta}{\sqrt{\sin^2 \frac{\theta_0}{2} - \sin^2 \frac{\theta}{2}}}$$

$$k = \sin \frac{\theta_0}{2} : \quad T = \frac{2}{k} \sqrt{\frac{l}{g}} \int_0^{\theta_0} \frac{d\theta}{\sqrt{1 - \frac{1}{k^2} \sin^2 \frac{\theta}{2}}}$$



Basic Concepts

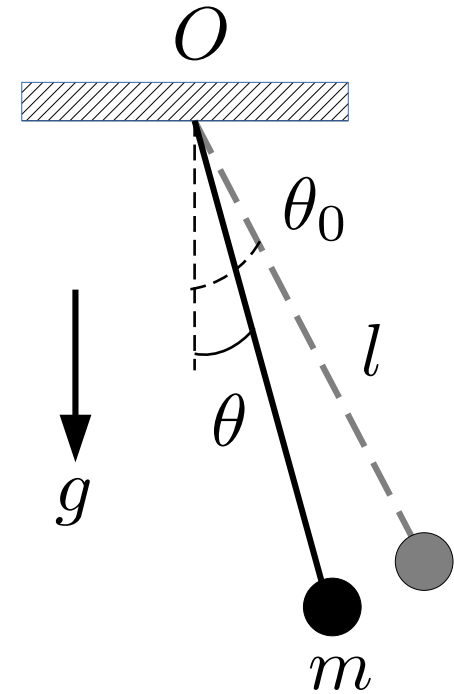
- A frictionless pendulum

$$T = \frac{2}{k} \sqrt{\frac{l}{g}} \int_0^{\theta_0} \frac{d\theta}{\sqrt{1 - \frac{1}{k^2} \sin^2 \frac{\theta}{2}}}$$

$$k \sin \alpha = \sin \frac{\theta}{2}, \quad 0 \leq \sin \frac{\theta}{2} \leq \sin \frac{\theta_0}{2} \Rightarrow 0 \leq \alpha \leq \pi/2$$

$$T = 4 \sqrt{\frac{l}{g}} \int_0^{\pi/2} \frac{d\alpha}{\sqrt{1 - k^2 \sin^2 \alpha}} = 4 \sqrt{\frac{l}{g}} F(k, \frac{\pi}{2})$$

$F(k, \frac{\pi}{2})$: Complete elliptic integral of the first kind

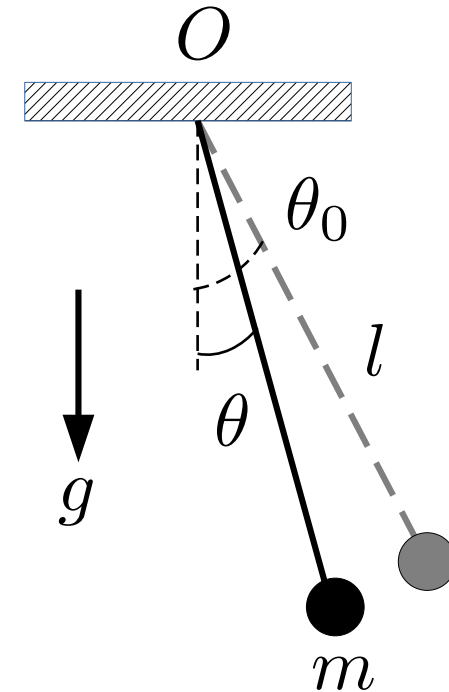


Basic Concepts

- A frictionless pendulum

$$T = 4\sqrt{\frac{l}{g}} \int_0^{\pi/2} \frac{d\alpha}{\sqrt{1 - k^2 \sin^2 \alpha}} = 4\sqrt{\frac{l}{g}} F\left(k, \frac{\pi}{2}\right)$$

Amplitude, θ_0	$k = \sin\left(\frac{\theta_0}{2}\right)$	$F\left(k, \frac{\pi}{2}\right)$	Period, T
0°	0	$1.5708 = \pi/2$	T_0
10°	0.0872	1.5738	$1.0019 T_0$
45°	0.3827	1.6336	$1.0400 T_0$
90°	0.7071	1.8541	$1.1804 T_0$
135°	0.9234	2.4003	$1.5281 T_0$
178°	0.99985	5.4349	$3.5236 T_0$
179°	0.99996	5.2660	$4.6002 T_0$
180°	1	∞	∞



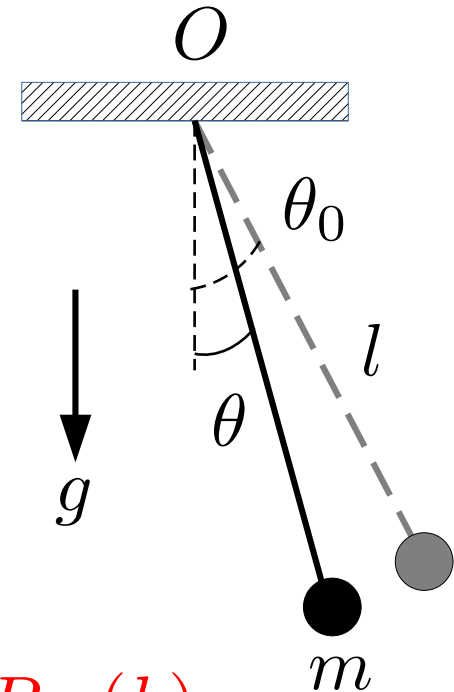
Basic Concepts

- A frictionless pendulum

$$T = 4\sqrt{\frac{l}{g}} \int_0^{\pi/2} \frac{d\alpha}{\sqrt{1 - k^2 \sin^2 \alpha}} = 4\sqrt{\frac{l}{g}} F\left(k, \frac{\pi}{2}\right)$$

$$F\left(k, \frac{\pi}{2}\right) = \frac{\pi}{2} \sum_0^{\infty} \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n} = \frac{\pi}{2} \sum_0^N \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n} + R_N(k)$$

~~$R_N(k)$: Truncation Error~~

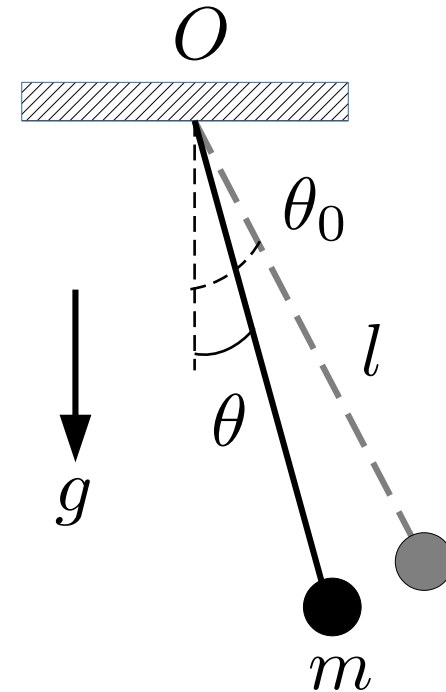


Basic Concepts

- A frictionless pendulum

$$T = 2\pi \sqrt{\frac{l}{g}} \sum_0^{\infty} \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n} = T_0 \sum_0^{\infty} \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n}$$

$$\sum_0^{\infty} \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n} = \frac{T}{T_0} \Rightarrow k = ? \text{ or } \theta_0 = ?$$



However, such an approach would be very inefficient due to two reasons:

- We are confronted with the **impossibility of finding analytically the roots** of a polynomial of order $N > 4$.
- At which value should we **truncate the power series**?

Basic Concepts

- Classify the occurring errors based on the structure every numerical routine.

- * **Input-Error**

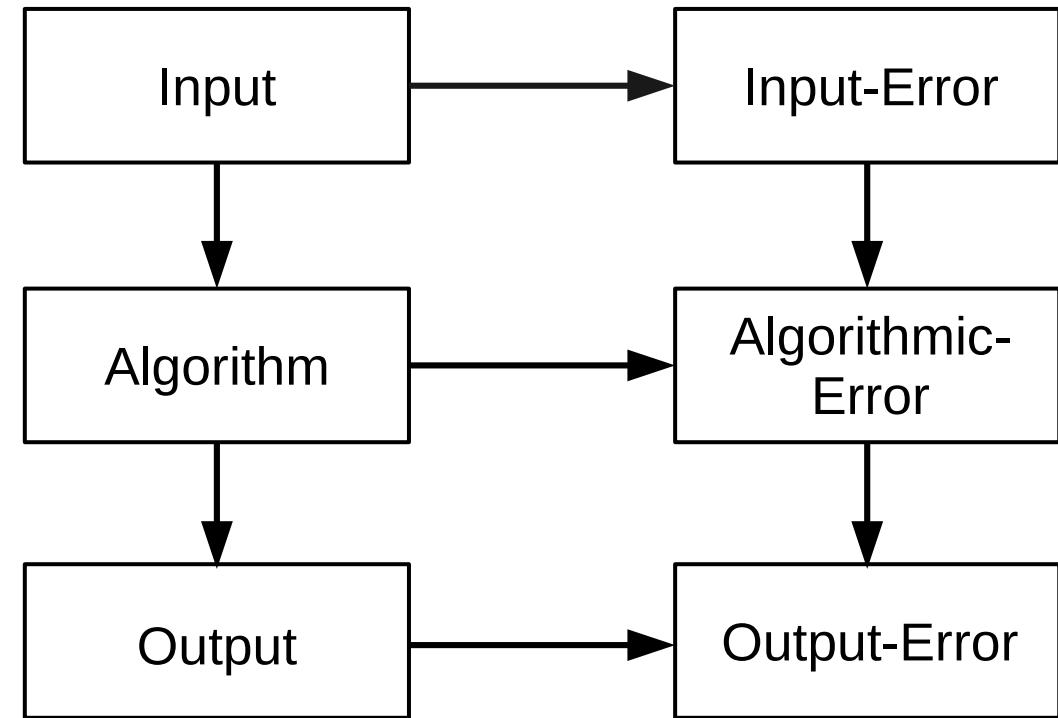
- Roundoff errors in input data

- * **Algorithmic-Error**

- Roundoff errors during evaluation
- Methodological errors due to mathematical approximations

- * **Output-Error**

- Roundoff errors



Basic Concepts

- Roundoff errors

Since every number is stored in a **computer using a finite number of digits**, we have to **truncate** every non-terminating number at some point.

$$\frac{2}{3} = 0.6666666666 \dots \xrightarrow{\text{stored}} \overbrace{0.6666666667}^{\text{Ten digits}} \text{ truncated}$$

floating-point form $\text{FL}\left(\frac{2}{3}\right) = 0.6666666667$

This has the consequence: $\text{FL}(\sqrt{3})\text{FL}(\sqrt{3}) \neq \text{FL}(\sqrt{3}\sqrt{3}) = 3$

Basic Concepts

- Roundoff errors

* Absolute error

$$\epsilon_a = |y - \bar{y}|$$

True value y
Approximate value \bar{y}

* Relative error

$$\epsilon_r = \left| \frac{y - \bar{y}}{y} \right| = \frac{\epsilon_a}{|y|}$$

In most applications, the **relative error is more significant.**

	y	\bar{y}	ϵ_a	ϵ_r
(1)	0.1	0.09	0.01	0.1
(2)	1000.0	999.99	0.01	0.00001

Basic Concepts

- Roundoff errors

$$\begin{array}{ll} \text{True value} & y = 0.d_1d_2d_3 \cdots d_kd_{k+1} \cdots \times 10^n \\ \text{Approximate value} & \bar{y} = 0.d_1d_2d_3 \cdots d_k \times 10^n \end{array} \quad d_1 \neq 0$$

$$\epsilon_r = \left| \frac{0.d_1d_2d_3 \cdots d_kd_{k+1} \cdots \times 10^n - 0.d_1d_2d_3 \cdots d_k \times 10^n}{0.d_1d_2d_3 \cdots d_kd_{k+1} \cdots \times 10^n} \right|$$

$$\epsilon_r = \left| \frac{0.d_{k+1}d_{k+2} \cdots \times 10^{-k}}{0.d_1d_2d_3 \cdots d_kd_{k+1} \cdots} \right| = \left| \frac{0.d_{k+1}d_{k+2} \cdots}{0.d_1d_2d_3 \cdots} \right| 10^{-k} < \frac{1}{0.1} 10^{-k} = 10^{-k+1}$$

$$d_1 \neq 0 : 0.d_1d_2d_3 \cdots \geq 0.1, \quad 0.d_{k+1}d_{k+2} \cdots < 1$$

Basic Concepts

- Roundoff errors

- * Particular care is required

- when two nearly identical numbers are subtracted**

- or

- when a large number is divided by a small number.**

- increase in the relative error.

- * In such cases the **roundoff error will increase** dramatically.

- * We note that it might be necessary to avoid such operations in our aim to design an algorithm which is required to produce reasonable results.

Basic Concepts

- Roundoff errors

* The **machine-number** is smallest positive number which can be added to another number, such that a change in the result is observed.

$$\eta = \min_{\delta} \{ \delta > 0 \mid 1 + \delta > 1 \}$$

* A typical value for **double-precision** in FORTRAN or C is

$$\eta \approx 10^{-16}$$

Basic Concepts

- Methodological Errors

* A **methodological error** is introduced whenever a **complex mathematical expression is replaced by an approximate**, or we have to deal with expressions we cannot evaluate analytically.

$$\sum_0^{\infty} \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n} \approx \frac{\pi}{2} \sum_0^N \left[\frac{(2n)!}{2^{2n} (n!)^2} \right]^2 k^{2n}$$

$$f'(x_0) = \left. \frac{df}{dx} \right|_{x_0} = \lim_{h \rightarrow 0} \frac{f(x_0 + h) - f(x_0)}{h} \xrightarrow{\text{Finite difference}} \left. \frac{df}{dx} \right|_{x_0} \approx \frac{f(x_0 + h) - f(x_0)}{h}$$