Computational Physics

M. Reza Mozaffari

Physics Group, University of Qom

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Central Forces



M. Reza Mozaffari

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 $\begin{cases} F(\vec{r})\hat{e}_r = \mu \frac{\mathrm{d}^2 \vec{r}}{\mathrm{d}t^2} \\ \frac{\mathrm{d}^2 \vec{R}}{\mathrm{d}t^2} = 0 \end{cases} \qquad \begin{cases} \vec{r} = r\hat{e}_r, & \frac{\mathrm{d}\vec{r}}{\mathrm{d}t} = \dot{r}\hat{e}_r + r\dot{\phi}\hat{e}_\phi \\ \frac{\mathrm{d}^2 \vec{r}}{\mathrm{d}t^2} = (\ddot{r} - r\dot{\phi}^2)\hat{e}_r + (2\dot{r}\dot{\phi} + r\ddot{\phi})\hat{e}_\phi \end{cases}$ $\stackrel{\times r}{\Longrightarrow} \mu(2\dot{r}\dot{\phi} + r\ddot{\phi}) = 0 \Rightarrow \mu(2r\dot{r}\dot{\phi} + r^2\ddot{\phi}) = 0$ $\frac{\mathrm{d}}{\mathrm{d}t}(\mu r^2 \dot{\phi}) = 0$

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$$\begin{cases} \mu(\ddot{r} - r\dot{\phi}^2) = F(r) \\ \frac{\mathrm{d}}{\mathrm{d}t}(\mu r^2 \dot{\phi}) = 0 \end{cases} \begin{cases} \vec{L} = \mu \vec{r} \times \vec{v} = \mu r^2 \dot{\phi} \hat{z} = L_z = l \\ \frac{\mathrm{d}}{\mathrm{d}t}(\mu r^2 \dot{\phi}) = 0 \end{cases} \stackrel{\text{Conservation of}}{\text{angular momentum}} \Rightarrow l = \mu r^2 \dot{\phi}$$



 $\mu \ddot{r} = F(r) + \frac{l^2}{\mu r^3}$

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 $\begin{cases} \mu \ddot{r} = F(r) + \frac{l^2}{\mu r^3} \\ \dot{\phi} = \frac{l}{\mu r^2} \end{cases}$







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$$E[r(t)] = \frac{1}{2}\mu\dot{r}^{2} + \frac{l^{2}}{2\mu r^{2}} + U(r)$$
$$\frac{dE}{dt} = 0$$
$$\frac{dE}{dt} = \mu\dot{r}\ddot{r} - \frac{l^{2}}{\mu r^{3}}\dot{r} + \frac{U(r)}{dr}\dot{r}$$
$$\frac{dE}{dt} = \dot{r}\left(\mu\ddot{r} - \frac{l^{2}}{\mu r^{3}} - F(r)\right) = 0$$
Conservation of total energy

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• Cross section of scattering



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• Cross section of scattering



- The number of particles scattered per unit time between θ and $\theta + d\theta$ is equal to the number incident particles per unit time between b and b + db.
- For incident flux Φ , the number of particles scattered into the solid angle $d\Omega = 2\pi \sin \theta d\theta$ per unit time is given by

$$N\mathrm{d}\Omega = N2\pi\sin\theta\mathrm{d}\theta = 2\pi b\mathrm{d}b\Phi$$

$$\frac{N}{\Phi} = \frac{b}{\sin\theta} \left| \frac{\mathrm{d}b}{\mathrm{d}\theta} \right|$$

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$$\begin{cases} \frac{N}{\Phi} = \frac{b}{\sin\theta} \left| \frac{\mathrm{d}b}{\mathrm{d}\theta} \right| \\ \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega} = \frac{N}{\Phi} \end{cases} \Longrightarrow \frac{\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega} = \frac{b}{\sin\theta} \left| \frac{\mathrm{d}b}{\mathrm{d}\theta} \right| \end{cases}$$

For classical Coulomb scattering,

$$U(r) = \frac{k}{r}$$

particle follows hyperbolic trajectory. In this case, a straightforward calculation obtains the Rutherford formula:

$$\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega} = \frac{b}{\sin\theta} \left| \frac{\mathrm{d}b}{\mathrm{d}\theta} \right| = \frac{k^2}{16E^2} \frac{1}{\sin^4\theta/2}$$

• Cross section of scattering



$$\frac{1}{2}m\dot{r}^{2} + \frac{l^{2}}{2mr^{2}} + U(r) = E = \frac{1}{2}mv_{0}^{2}$$
$$\frac{1}{2}m\dot{r}^{2} = E - \frac{l^{2}}{2mr^{2}} - U(r)$$
$$\frac{1}{2}m\dot{r}^{2} = E\left(1 - \frac{l^{2}}{2Emr^{2}} - \frac{U(r)}{E}\right)$$
$$\frac{1}{2}m\dot{r}^{2} = \frac{1}{2}mv_{0}^{2}\left(1 - \frac{b^{2}}{r^{2}} - \frac{U(r)}{E}\right)$$
$$\dot{r}^{2} = v_{0}^{2}\left(1 - \frac{b^{2}}{r^{2}} - \frac{U(r)}{E}\right)$$
$$\dot{r} = \pm v_{0}\sqrt{1 - \frac{b^{2}}{r^{2}} - \frac{U(r)}{E}}$$

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$$\dot{r} = \frac{\mathrm{d}r}{\mathrm{d}t} = \pm v_0 \sqrt{1 - \frac{b^2}{r^2} - \frac{U(r)}{E}}$$
$$\dot{\phi} = \frac{\mathrm{d}\phi}{\mathrm{d}t} = \frac{l}{mr^2}$$
$$\frac{\dot{r}}{\dot{\phi}} = \frac{\mathrm{d}r}{\mathrm{d}\phi} = \pm \frac{mv_0 r^2 \sqrt{1 - \frac{b^2}{r^2} - \frac{U(r)}{E}}}{l}$$
$$\frac{\mathrm{d}r}{\mathrm{d}\phi} = \pm \frac{r^2 \sqrt{1 - \frac{b^2}{r^2} - \frac{U(r)}{E}}}{b}$$

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$$\theta = \pi - 2b \int_{r_m}^{\infty} \frac{\mathrm{d}r}{r^2 \sqrt{1 - \frac{b^2}{r^2} - \frac{U(r)}{E}}}$$

$$\forall b \to r_m, \theta, \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega}$$

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